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Atmospheric Infrared Sounder

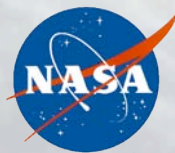
Improvements For V6 To Handle Channel Frequency Shifts

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April 17, 2008

**AIRS Science Team Meeting
April 15–17, 2008, Caltech**

Handling Frequency Shifts In V6



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Introduction

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- In this talk I will describe work in progress to enable V6 to account for the small, time-dependent frequency shifts of the AIRS IR channels
 - *The existing L1B algorithm to measure the shifts will be improved*
 - *The dynamic frequencies will be used in Level 2*
 - *A Level 1C product will be generated which will consist of cleaned-up spectra optionally resampled onto a fixed frequency grid*
- Many people are contributing to this effort, including:
 - *George Aumann* *Larrabee Strow*
 - *Yibo Jiang* *Scott Hannon*
 - *Evan Manning*
 - *Margie Weiler*



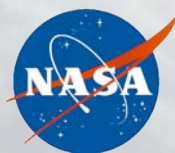
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Outline

- The need for better handling of frequency shifts
- Requirements for V6
- Preliminary design and algorithm descriptions
 - *The four components of the shifts*
 - *Dynamic determination of frequency shifts*
 - *Noisy channels detection and monitoring*
 - *Cleaning up noisy spectra*
 - *Shifting to a fixed frequency grid*
- Summary of planned software changes



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Why do we need to handle the extremely small frequency shifts?

- We have always known that AIRS channel frequencies vary slightly with time, due primarily to small changes in temperature gradients in the spectrometer
- Using a fixed frequency set in L2 works fine for meeting our 1K / 1 km primary requirement, weather prediction, and many other purposes
- For climate studies, where we want to measure trends at the 10 mK/yr level, we need to properly account for the shifts in both L1 and L2
 - *Existing L1B code attempts to measure the shifts, but the results are noisy and V5 makes no use of the results*
- Thus we have three new top-level requirements for V6
 - *Measure frequency shifts as accurately as we can*
 - *Provide a prescription (and possibly a product) for resampling radiances to a truly fixed frequency grid (Level 1C)*
 - *Account for the dynamic frequency shifts during retrievals*



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Detailed Requirements (1 of 2)

- **L1B—**
 - *Determine and record instantaneous frequencies of all channels*
 - *Provide a list of noisy channels for use in Level 1C*
- **L2—**Modify the RTA to make use of the actual frequencies that were determined in L1B



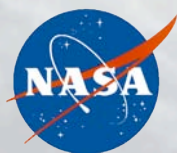
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Detailed Requirements (2 of 2)

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- **L1C—(new)**
 - Define a set of fixed channel frequencies, including pseudo-channels to fill gaps in existing spectral coverage
 - Generate (or provide prescription for) Level 1C products
 - “Clean” product
 - Replace radiances of noisy channels and supply radiances of pseudo channels using best available radiances from correlated good channels
 - Do not disturb L1B radiances of good channels
 - Supply information specifying whether radiance is NIST traceable (good channels) or not (noisy and pseudo channels)
 - Resampled product
 - Resample the “clean” spectra onto the fixed frequency grid



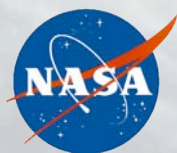
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Possible L1C Output Options

- Just provide a routine for users to generate their own “clean” and (optionally) resampled spectra
- Have Level 1C products (“clean” and/or resampled) produced at the GES DISC, but only by request for user-specified granules
- Routinely output a full Level 1C product (cleaned and resampled calibrated radiances)

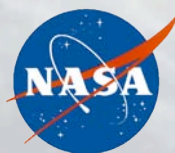


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Frequency shifts and their dynamic determination



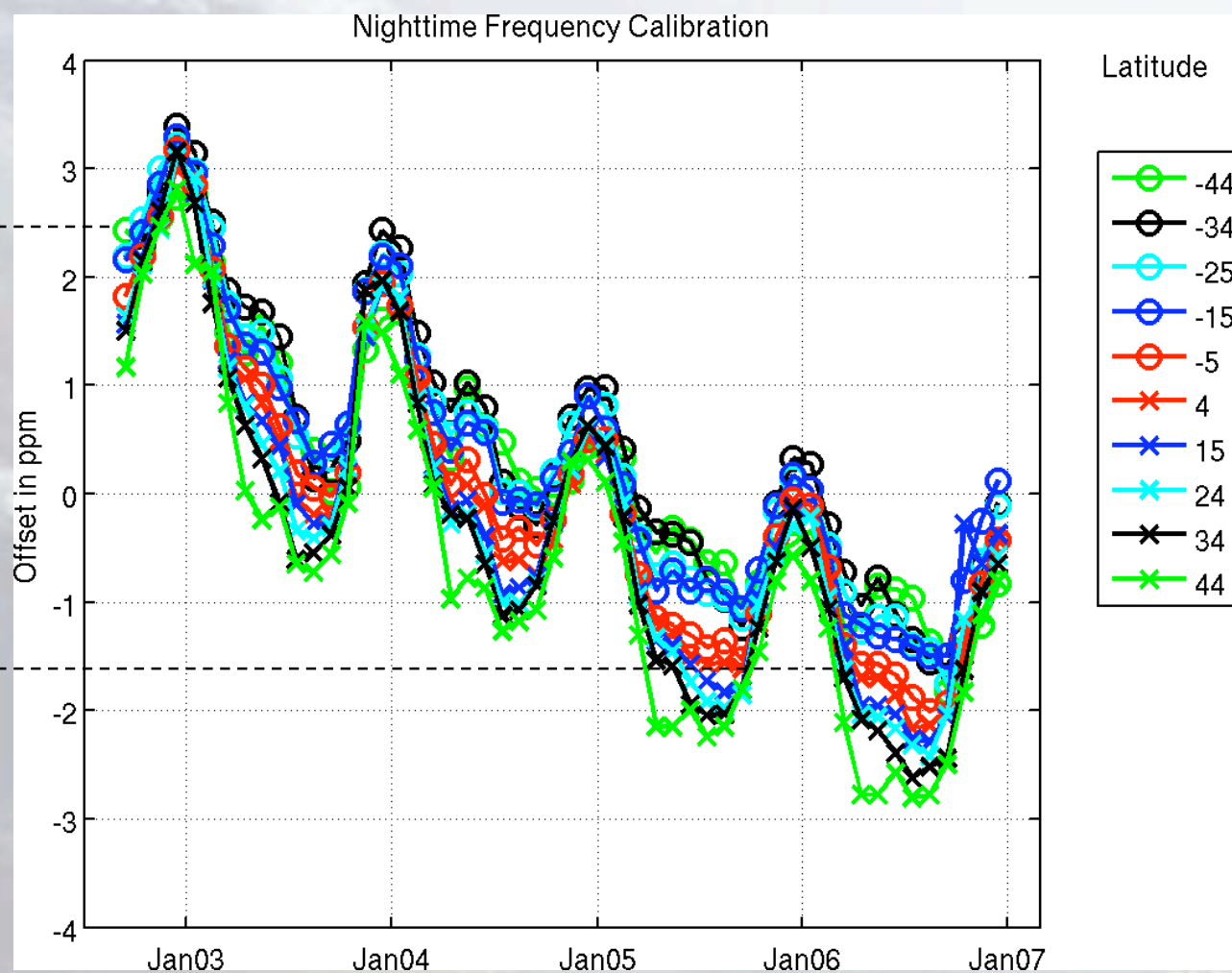
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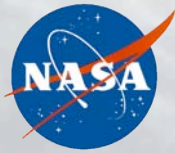
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Observed Frequency Shifts

$< 1 \text{ ppmf/yr}$





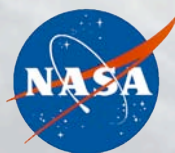
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Frequency Shifts

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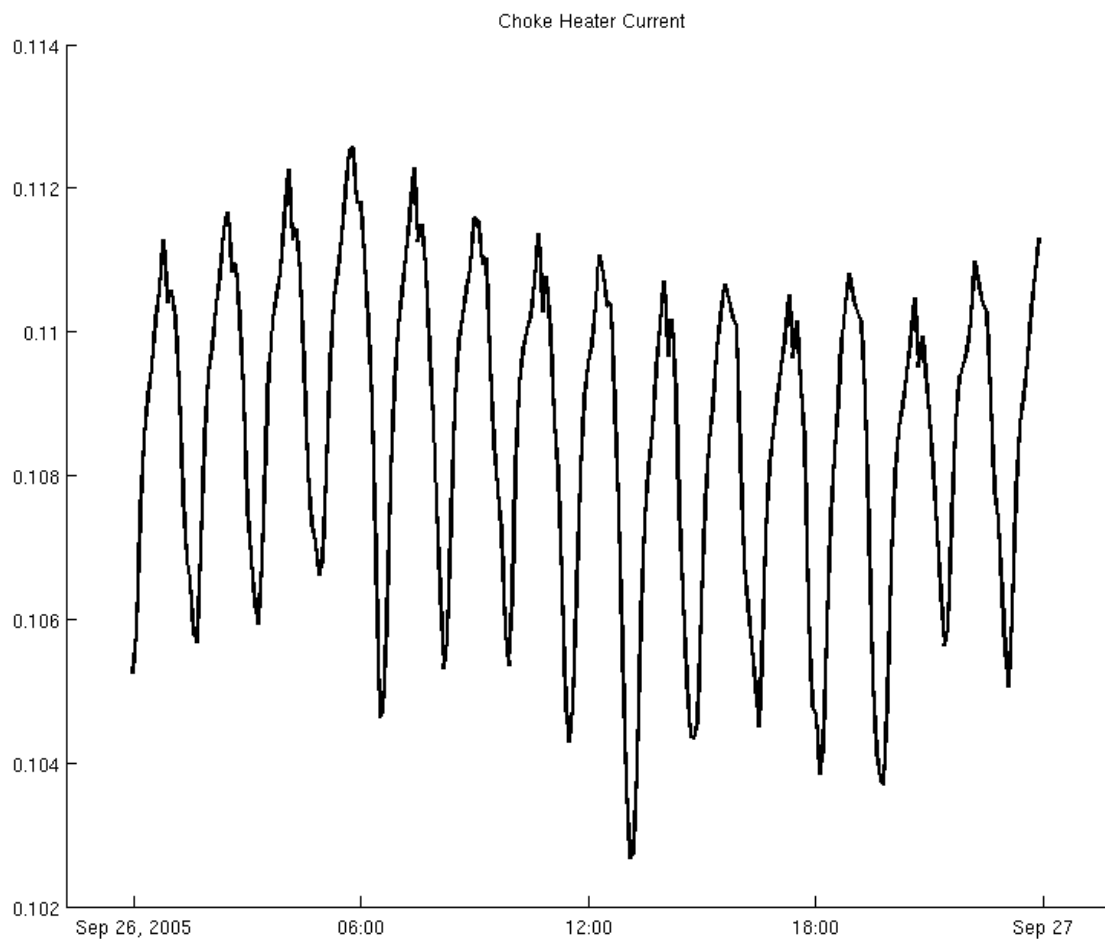
- **Results from CO₂ and H₂O channels are very similar which implies that all the detector modules are shifting together**
- **Latitude is used in the previous chart as a rough proxy for orbital position**
 - ***At least two methods for predicting the orbital shift are being investigated***
 - 1) Using time and orbital phase information to determine the time since entrance into or exit from the earth's shadow
 - 2) Using the current of the choke point heater, which tries to maintain the spectrometer at constant temperature by controlling a heater on the second stage radiator
- **Strow and Hannon showed that the seasonal oscillation with peak-to-peak amplitude 3 ppmf tracks the solar beta angle**
- **There is a secular change of approximately 1 ppmf/yr**
- **There is also 24-hour cycle in spectrometer temperatures which is captured by the choke point heater current**

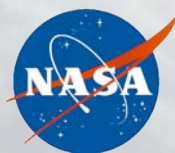


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Choke Point Heater Current Typical Day





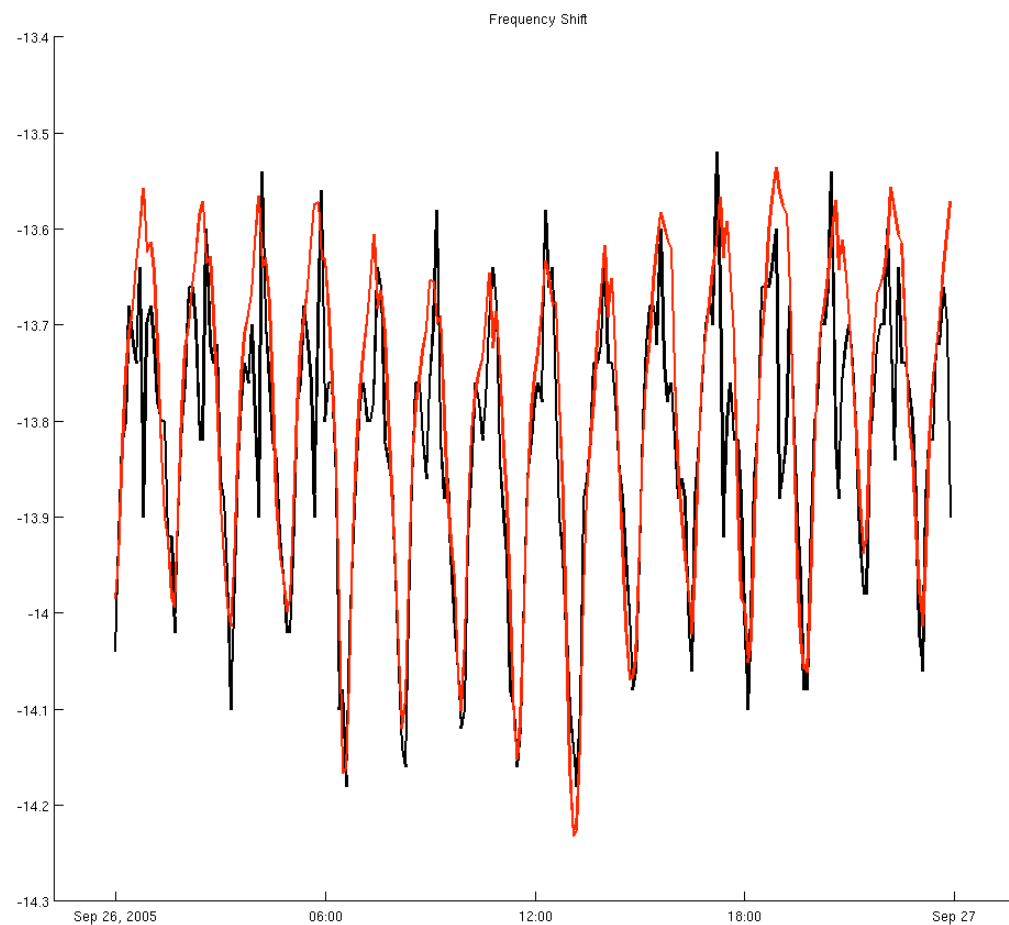
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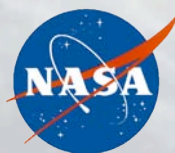
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Frequency Shift (Actual and Predicted)

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- **Black—Actual**
(measured by
Hannon)
- **Red—Predicted**
(from observed
choke point
heater current)





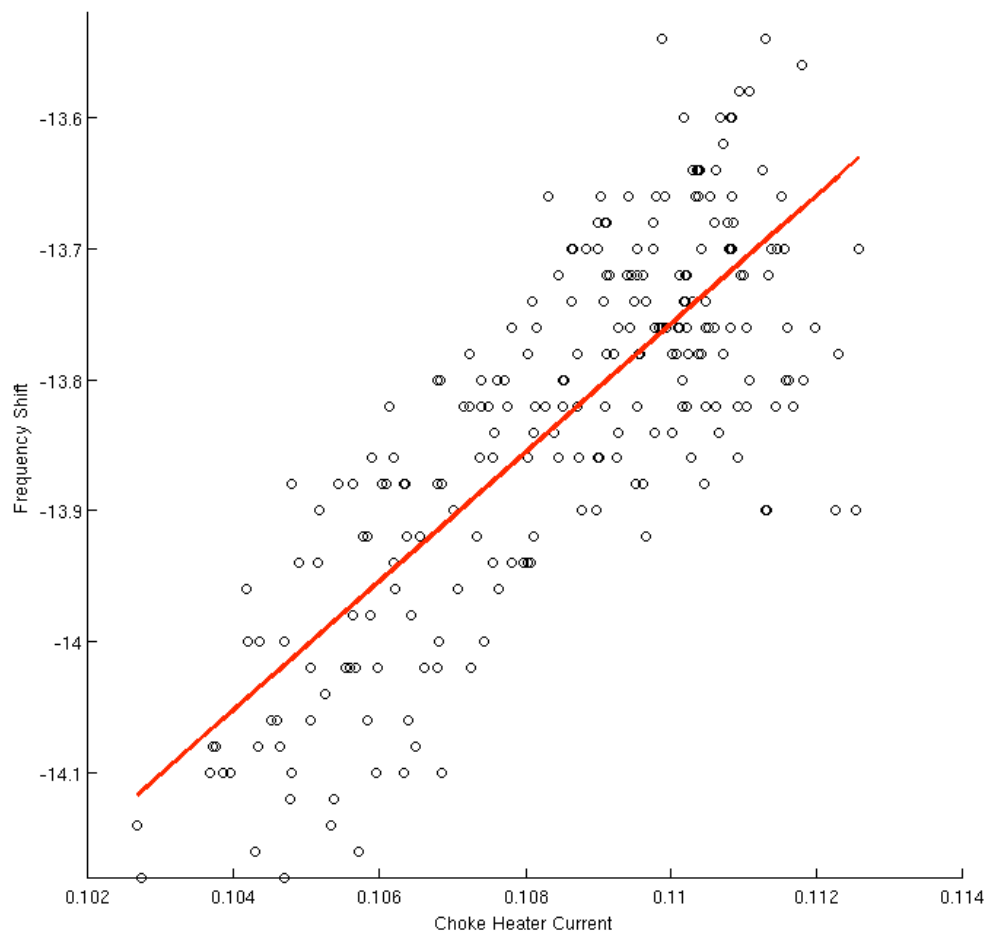
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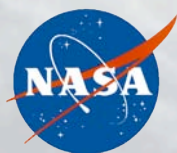
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Heater Current vs. Frequency Shift

- **Correlation coefficient is 0.78**



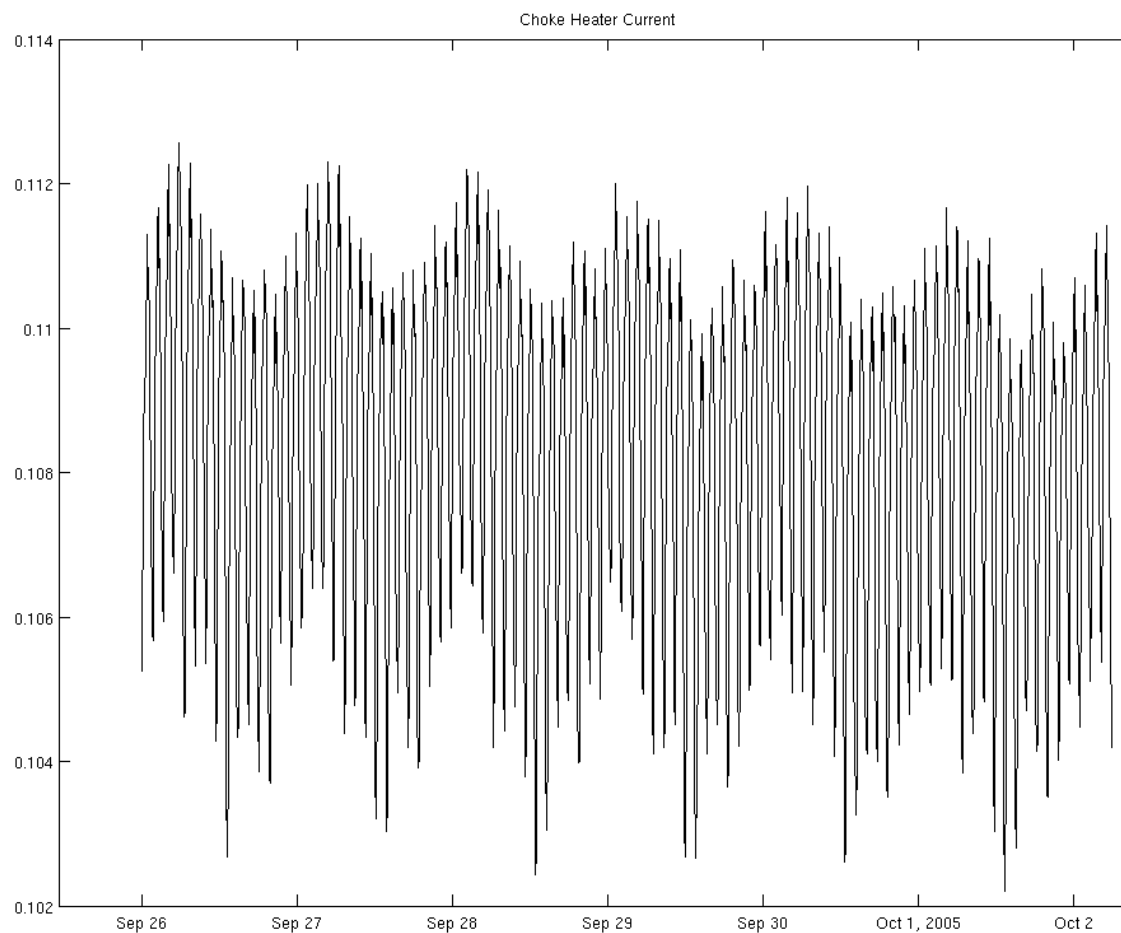


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Daily Cycle As Seen In Choke Heater Current

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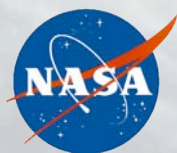
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Dynamic Frequency Determination

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- The instantaneous frequencies can be determined from three or four predictors known at run time
 - ***Orbital component***
 - Choke point heater current or
 - Time since terminator crossing
 - ***Daily component***
 - Choke point heater current or
 - UTC
 - ***Seasonal component***
 - Solar beta angle
 - ***Secular component***
 - Date



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Detection and Monitoring of Noisy Channels



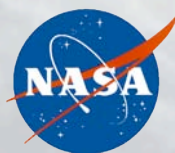
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Noisy Channel Detection Using PCA

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- **Once per mission, train on a set of 48 simulated AIRS clear spectra**
- **At a TBD frequency, use a test spectrum (actually observed) determined to be clear from a sea surface temperature comparison test**
 - *Use principal component analysis to calculate a reconstructed spectrum from the test spectrum*
 - *Analyze the brightness temperature difference between the raw test spectrum and the reconstructed*
 - Mark a channel “replaceable” if the difference is large ($> 10K$) or if the difference divided by $(1 + NE\Delta T)$ exceeds 5
 - *Generate a known “replaceable” channels list by concatenating the list of channels declared dead before launch to ones found to be bad by the PCA analysis*

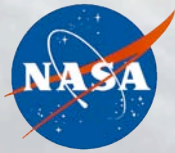


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Replacement Of Noisy Channel Radiances



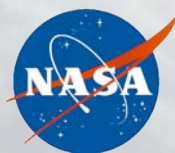
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Radiance Replacement (“Clean”)

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- The goal of spectrum clean-up (radiance replacement) is to permit interpolation and resampling without fear of introducing artifacts
- For each AIRS IR channel, a list of up to 10 possible replacement channels has been determined
 - *A channel can replace another channel if their brightness temperatures are expected to be close AND*
 - *If the channels are both window channels or both water channels or both ozone (etc.) AND*
 - *If the channels are close in frequency*
- For each spectrum, for each channel on the “replaceable” list, set the radiance to a TBD combination of channels from its replacement set
- This “cleaning” process will not disturb radiances of channels not marked for replacement



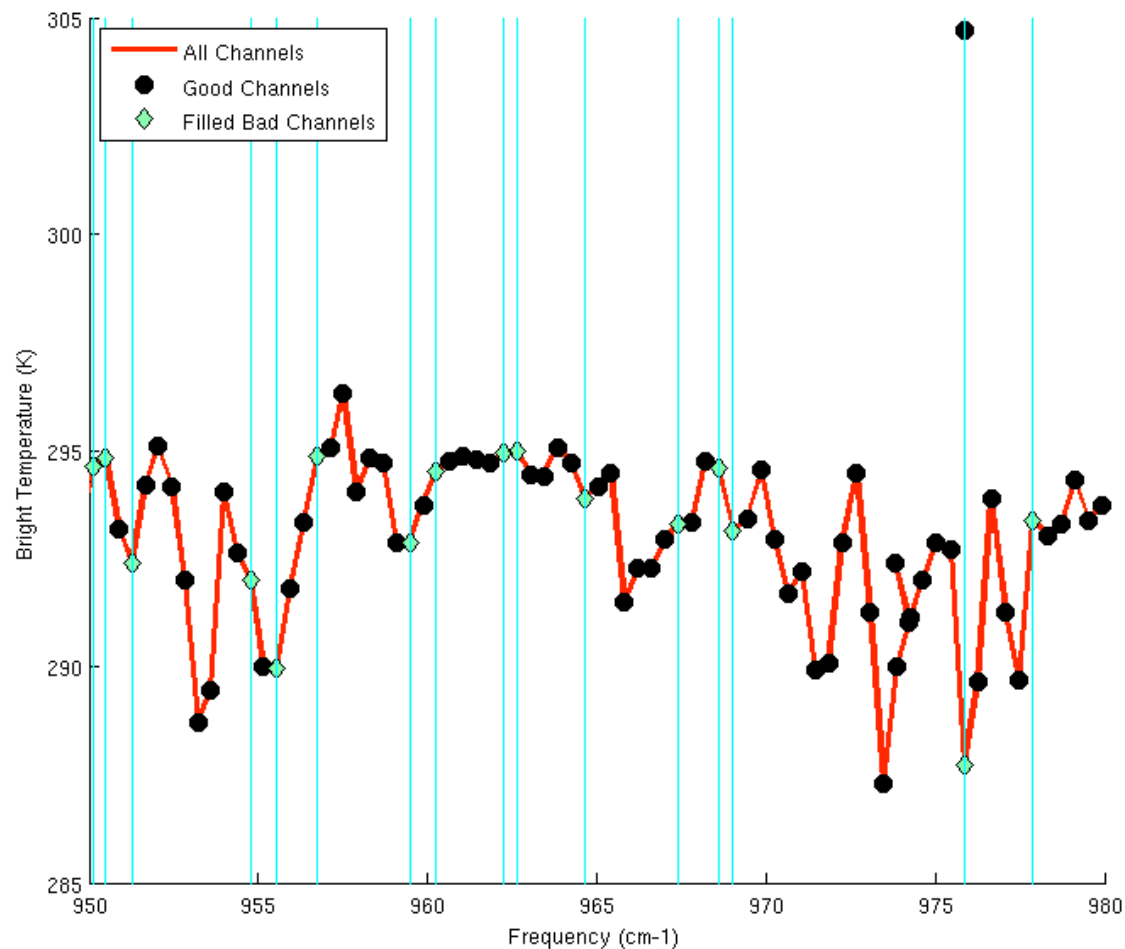
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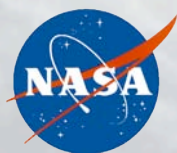
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Radiance Replacement Example

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- This range of frequencies is the worst case over the entire focal plane—it has an unusual concentration of noisy channels



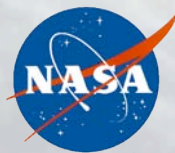


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Resampling



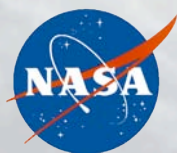
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Resampling

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- The cleaned-up spectra can be resampled to a fixed frequency grid (probably the nominal frequency set we now have) once the instantaneous frequency of each channel has been determined
- The interpolation technique now being studied is module-by-module spline interpolation



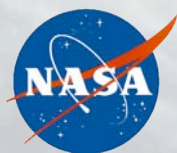
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Planned Software Changes (Summary)

- **Level 1B will be modified to calculate and store instantaneous channel frequencies**
- **Level 1B will be modified to perform the check for “replaceable” channels and store results**
- **The Level 1B output radiances will not change**
- **A new routine will be created to generate a Level 1C product**
 - *Perform the radiance replacement (“clean”)*
 - *Optionally resample to fixed frequency grid*
- **The RTA will be modified to use the instantaneous frequencies instead of the nominal set for retrievals**

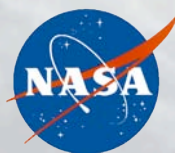


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• Backup



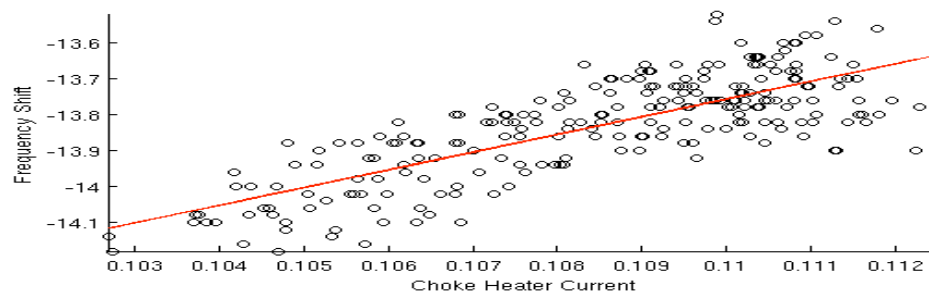
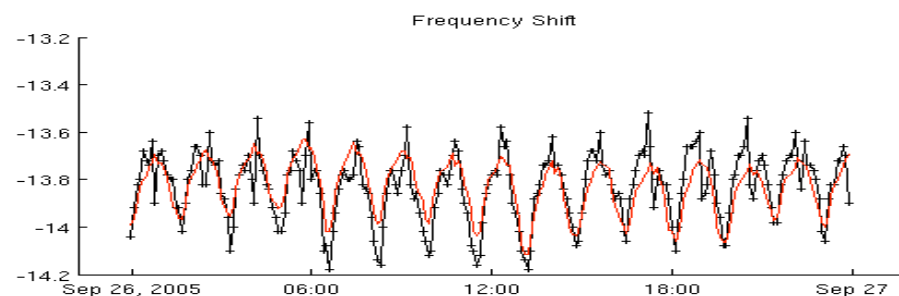
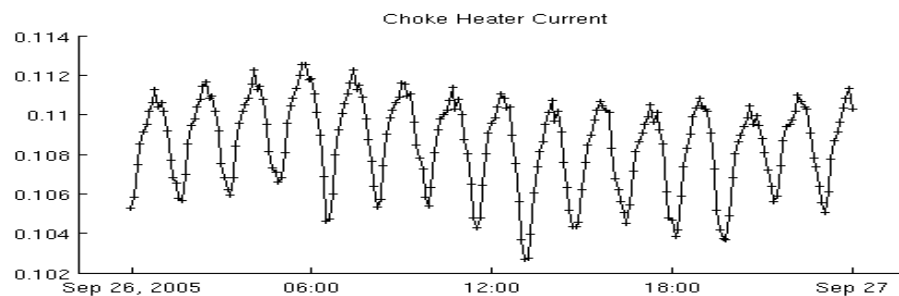
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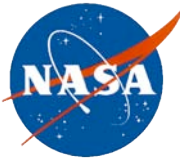
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- Choke point heater current for one day
- Black is measured frequency shift from Scott Hannon
- Red is frequency shift calculated from instantaneous value of choke point heater current
- Linear fit of choke point heater current vs. frequency shift
- Correlation is 0.78

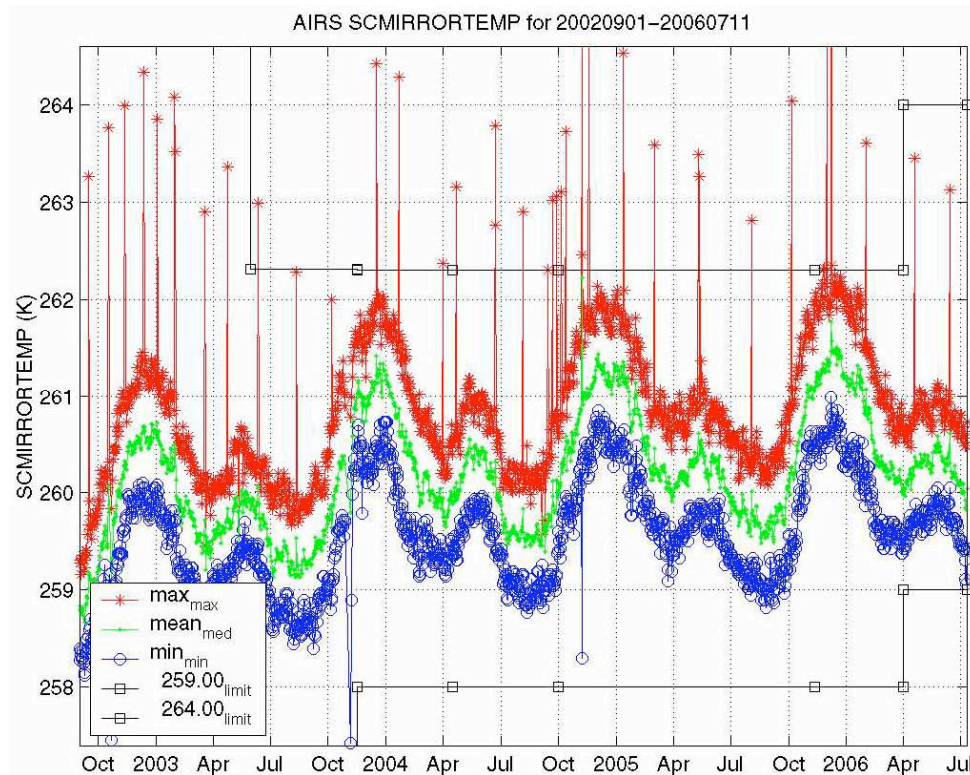
Choke Heater Current and Frequency Shift



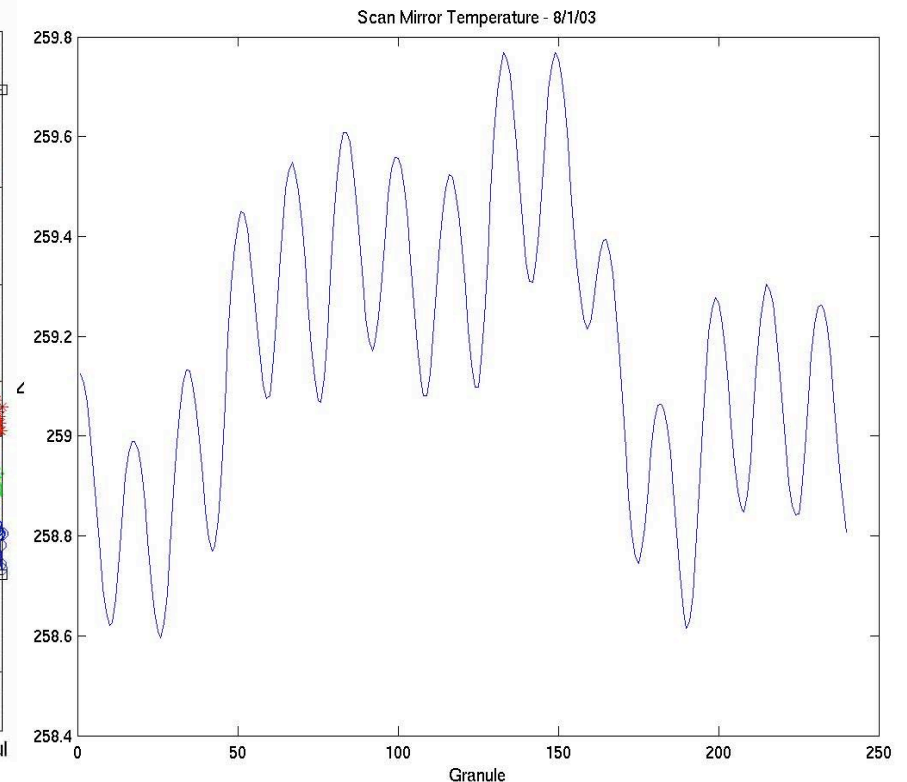


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AIRS Scan Mirror Temperature Varies with Year, Season, Orbit, and S/C Maneuvers

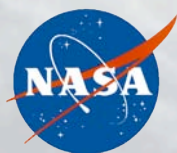


Temperature over mission



Temperature over 1 day

- **4 K spikes due to S/C maneuvers (mirror parked on OBC), 3K seasonal variation, 1K daily variation, .6 K orbital variation, as well as a upward trend of ~0.15 K per year.**



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Effect Of Clouds On Replacement

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